The ULB Binary-Star Modelling Project



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The ULB Binary-Star Modelling Project

Why binaries?

Many (most?) stars are in binaries

Accurate stellar parameters (L, R, M, T_{eff})

Exotic evolution

Pulsations! Ejections!

Accretion! Explosions!





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Observations:	M,MS,S (AGB)	Ba, CH, CEMP
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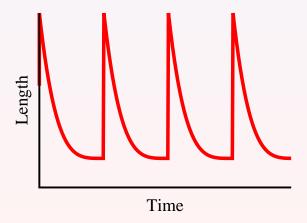
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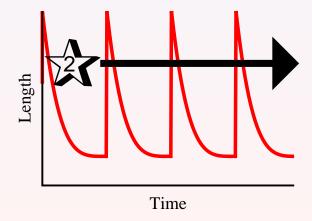
Scientific challenges

- ► Focus on low- and intermediate-mass binary stars
- ► AGB mass transfer : Ba, CH, CEMP, post-AGB



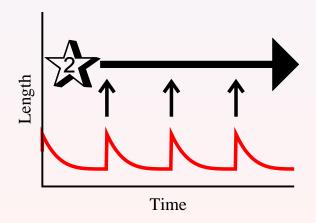
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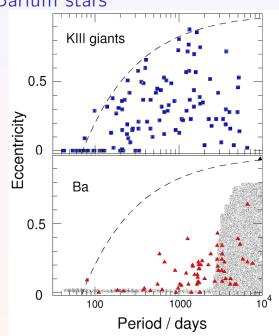


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e.g. Barium stars



STAREVOL (Lionel's code)

▶ Reliable, well-tested single-star evolution code

Derived from Kippenhahn's original code

- via Forestini etc.
- State-of-the-art input physics
- Detailed nucleosynthesis
- (S)AGB calculations, planetary accretion, rotational and gravity wave mixing etc.

Binary STAREVOL (Binstar? Binevol?)

- Complete rewrite!
 - FORTRAN 77→90, Modular
 - Numerical improvements
- Current status:
 - (most) Code reorganised
 - Evolves two stars at once
 - Basic mass transfer/orbital evolution
 - Working on the binary interaction physics (RLOF/wind/disc interaction)
 - Needs a better name!



Technological challenges

- ► Two stars→>twice as slow
- Parallel computing is the future
 - Quad-core CPU now
 - 8-thread CPU now
 - Open-MP parallelisation
- ► Graphics card: 100s of threads
 - ► Double precision now/soon
 - New scientific libraries
 - Charal







Final Remarks

- ▶ Binary-star code ready soon:
 - Tests new physics (e.g. wind/disc-interactions)
 - Self-consistent detailed AGB binary models!
 - Complements observational surveys
 - Harness new technology