

Barium Stars Revisited



Robert Izzard



and



Tyl Dermine

Université Libre de
Bruxelles

and

Ross Church

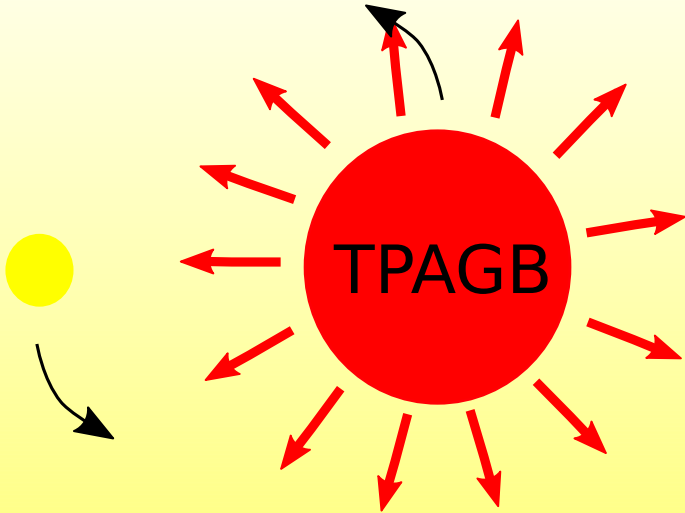
Lund Observatory



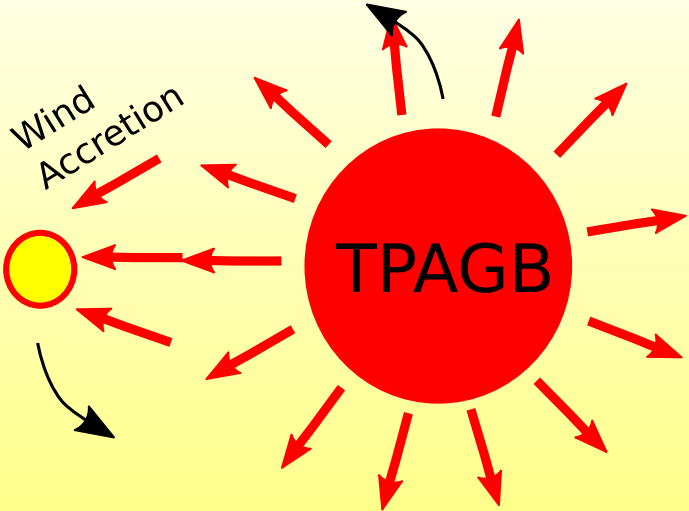
What is a barium star?



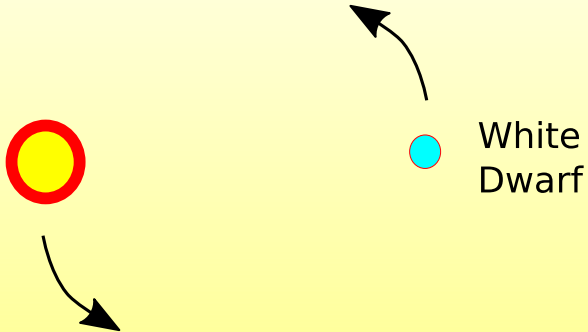
How to make a Ba star? A long time ago...



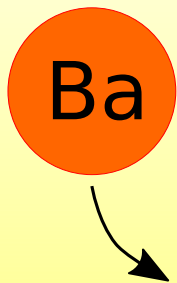
How to make a Ba star? A long time ago...



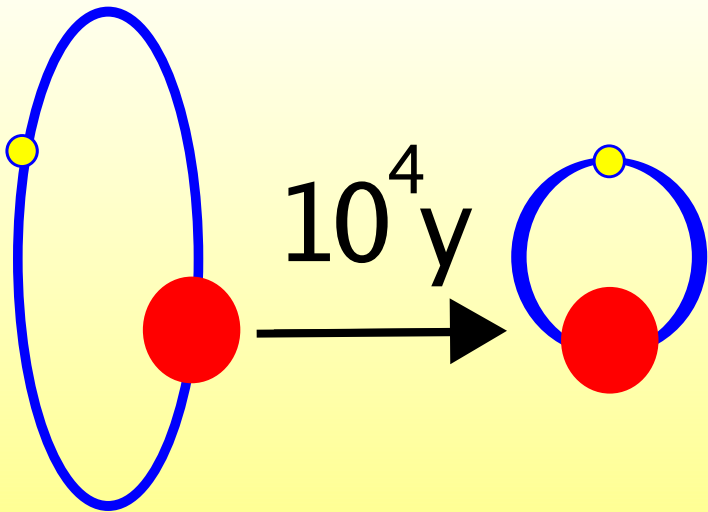
How to make a Ba star? A long time ago...



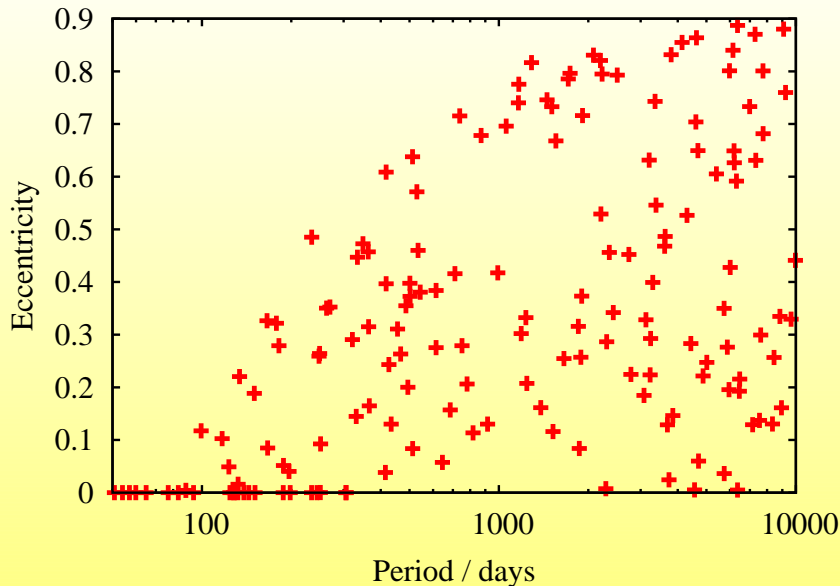
The barium star now



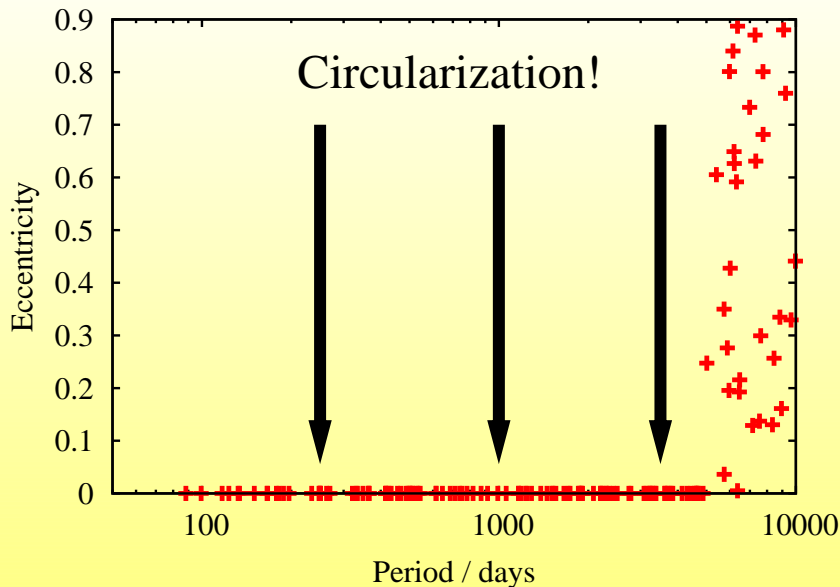
Tidal Circularization $\tau \sim (a/R)^8$



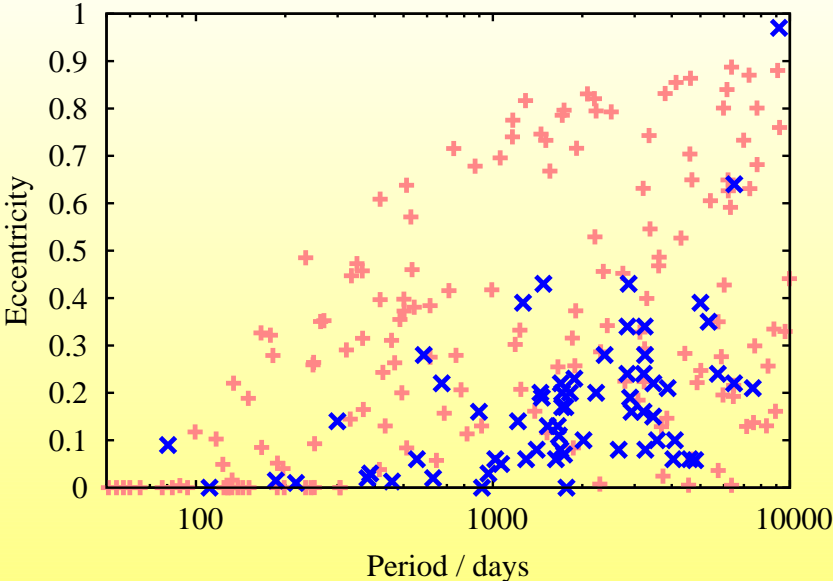
Observations: GK giants (Jorissen data)



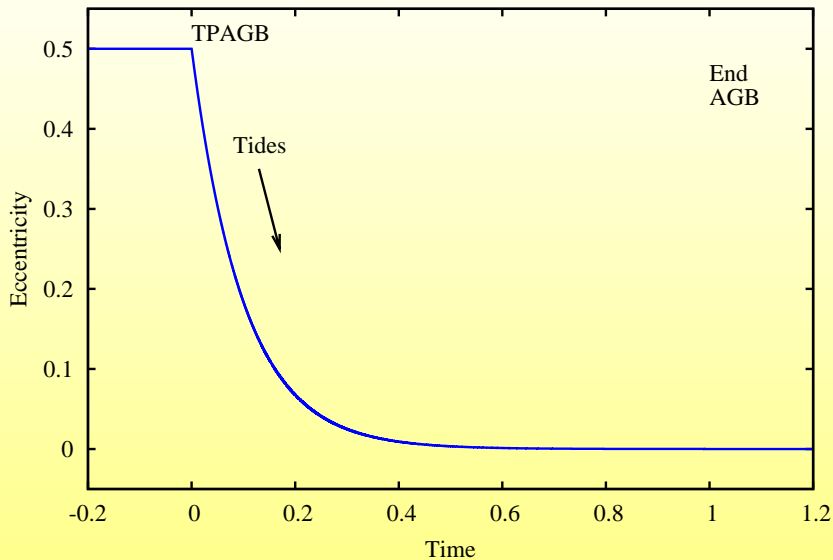
Expected result for Ba stars



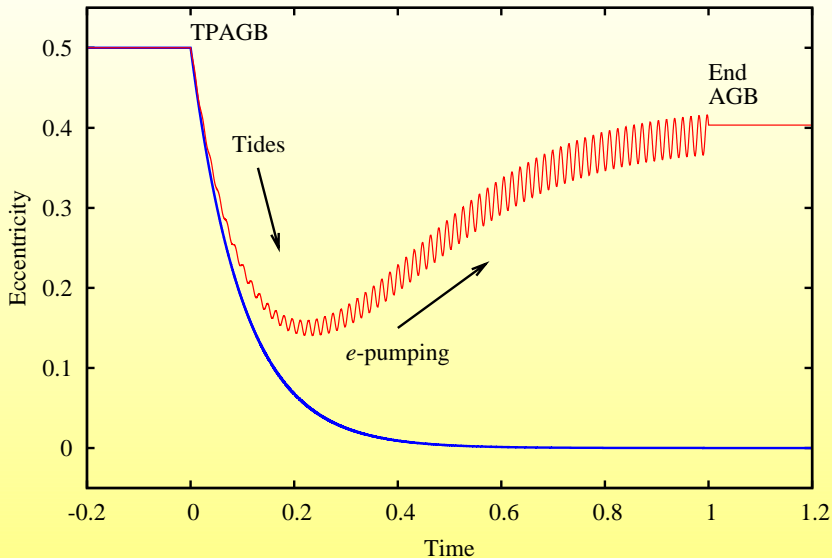
But it doesn't/mayn't happen!



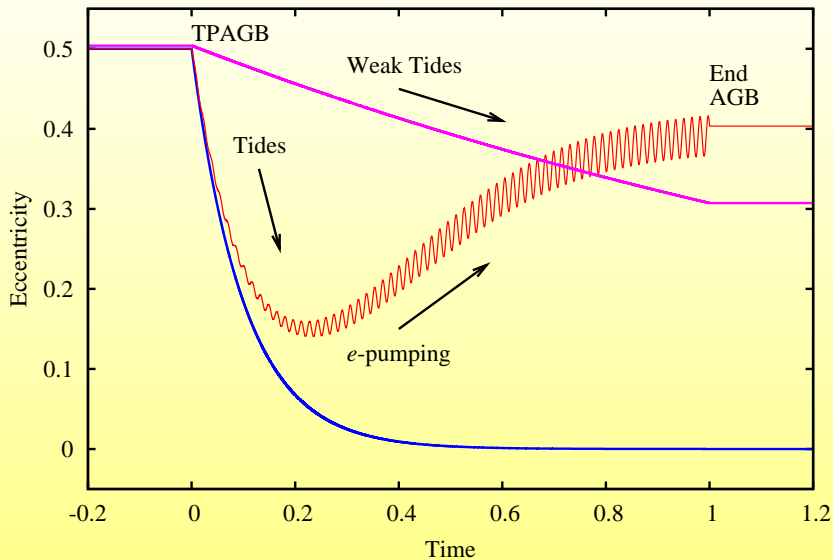
What is happening?



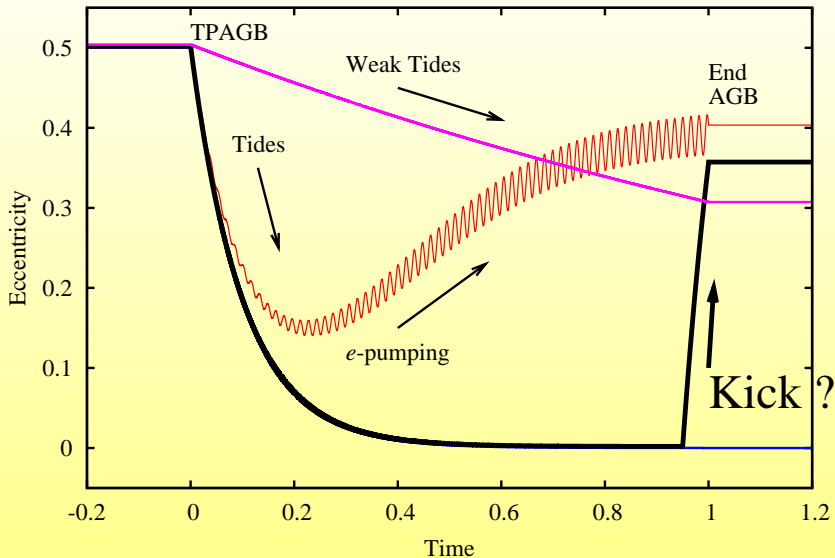
What is happening?



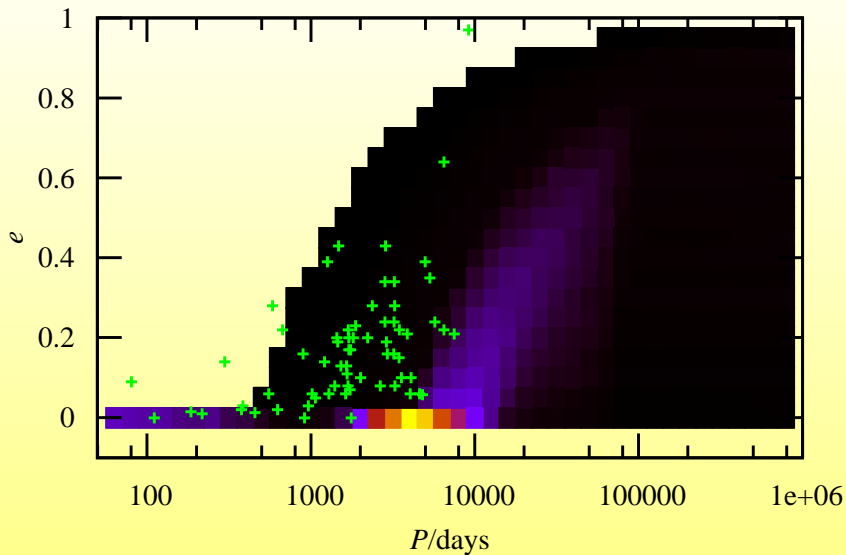
What is happening?



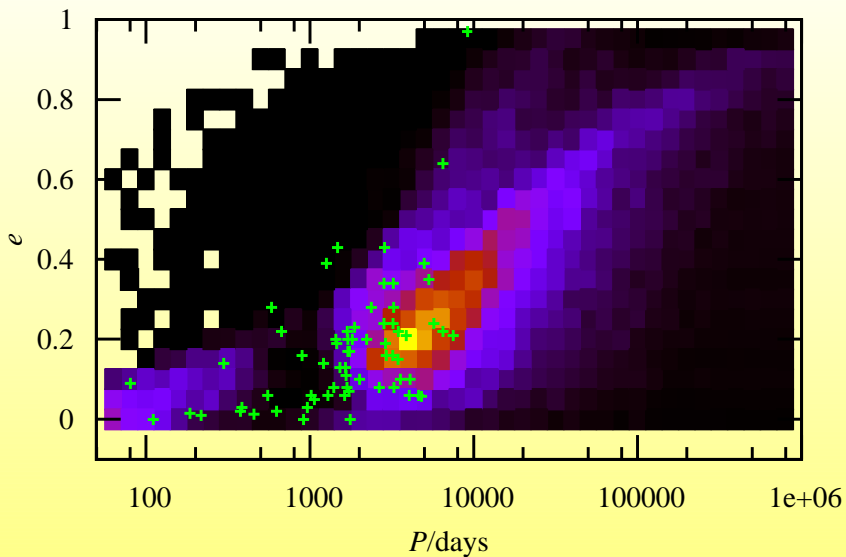
What is happening?



Canonical result

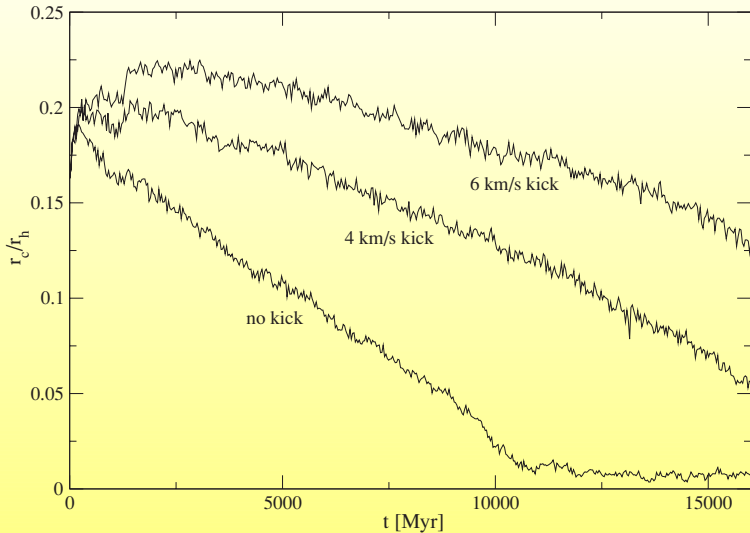


4km/s kick

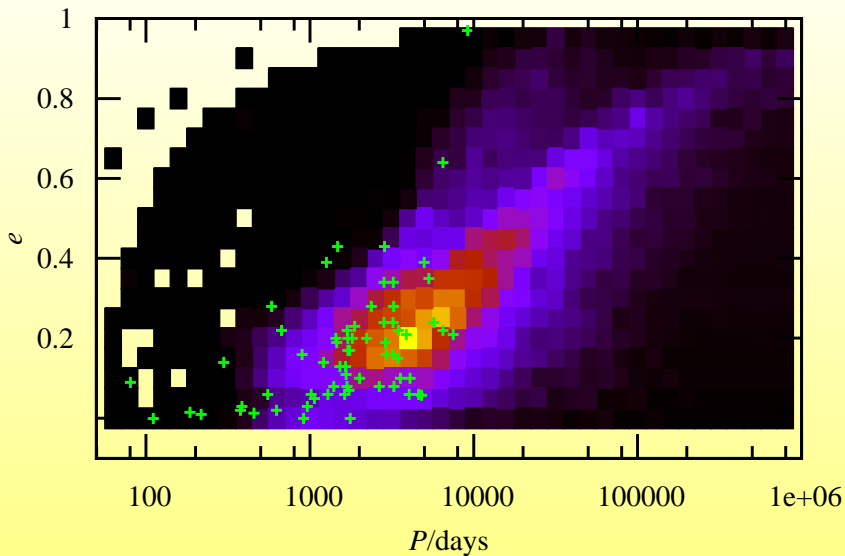


Independent evidence from GCs

Fregeau, Heyl etc.



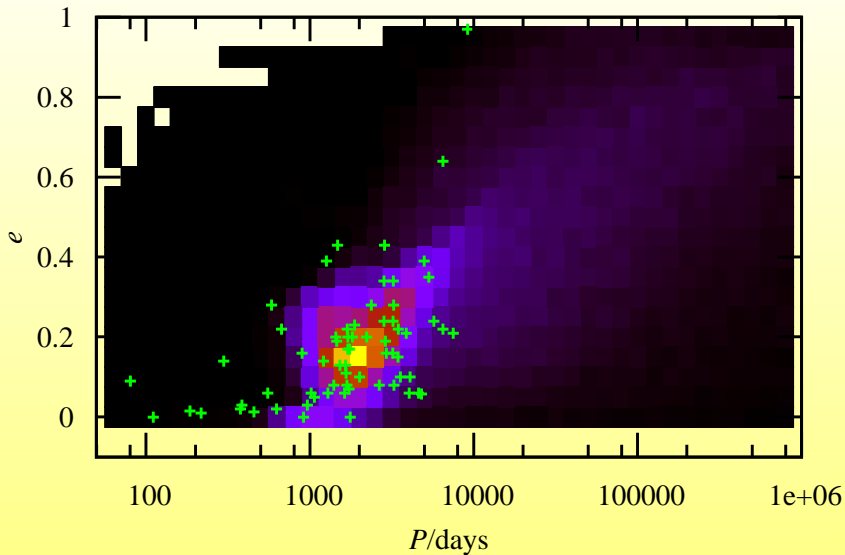
4km/s kick+efficient CE



Efficient common envelope ejection?

- ▶ *As if we have a clue?!*
- ▶ Envelope weakly bound
- ▶ Even with $\alpha = 1$ only a few % of envelope's recombination energy is required
- ▶ Nelemans/Tout γ prescription \rightarrow similar!
- ▶ Only for $P \lesssim 2000$ days

4km/s kick + efficient CE and $\dot{M}_{\text{acc}} \rightarrow \dot{J}$



Angular momentum?

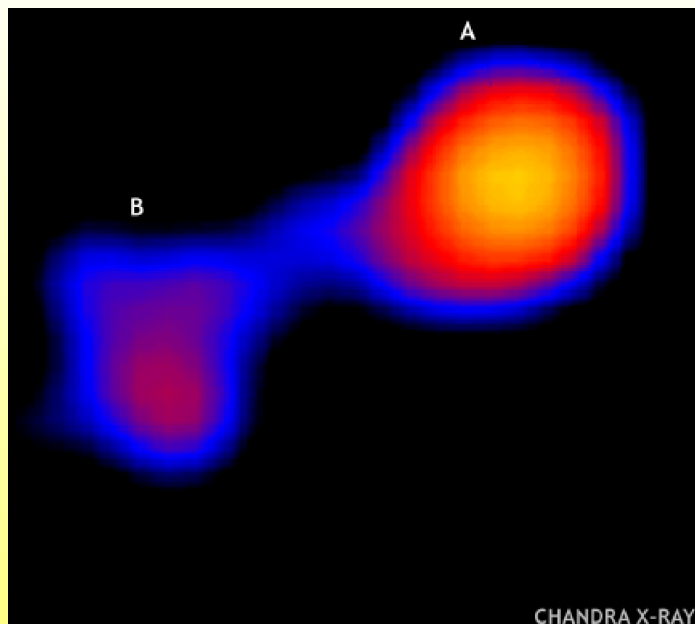
- ▶ Huang 1956:
 - ▶ *when the lost mass is collected by the secondary component, the major axis will decrease*

▶

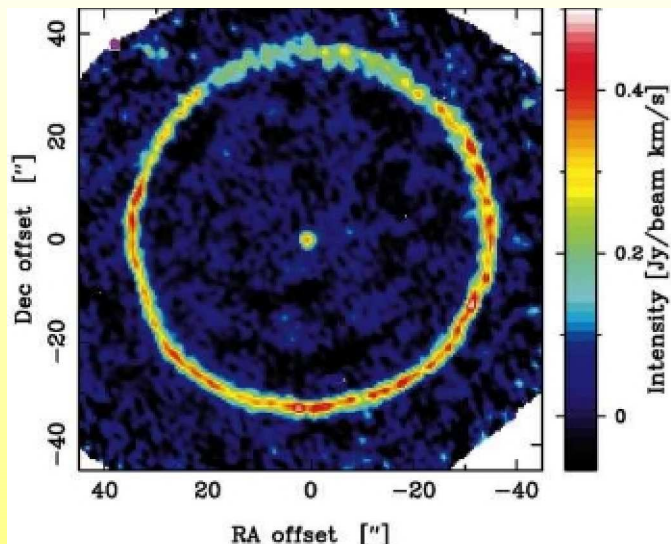
$$\dot{J}_{\text{orbit}} \propto \left(-|\dot{M}_{\text{lost}}| - \frac{M_2}{M_1} |\dot{M}_{\text{acc}}| \right)$$

- ▶ Second term **large** for efficient accretion!
- ▶ Only for wide systems $P \gtrsim 5000$ days
- ▶ Alternatively:
 - ▶ magnetic braking
 - ▶ disc mass loss?

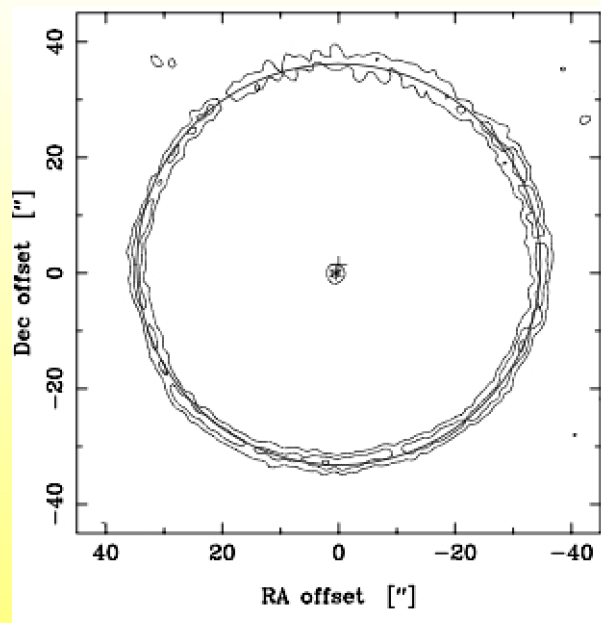
Efficient accretion: Mira



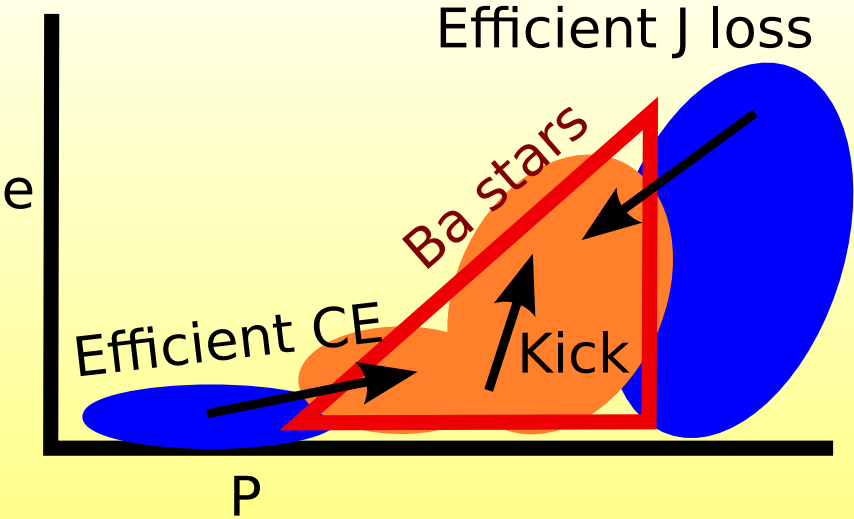
Kick Implications: PNe (TT Cygni in CO)



Kick Implications: PNe (TT Cygni in CO $v \sim 0.6 \text{ km s}^{-1}$)



Conclusion: How To Make Ba Stars...?



Big Brains→A better solution?

